

PLATO instrument performance

Anko Börner, Carsten Paproth

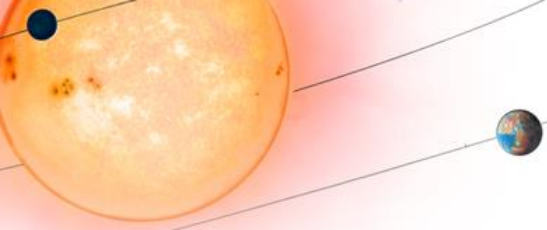
STESCI Workshop II

Milazzo, May 22nd – 25th 2017

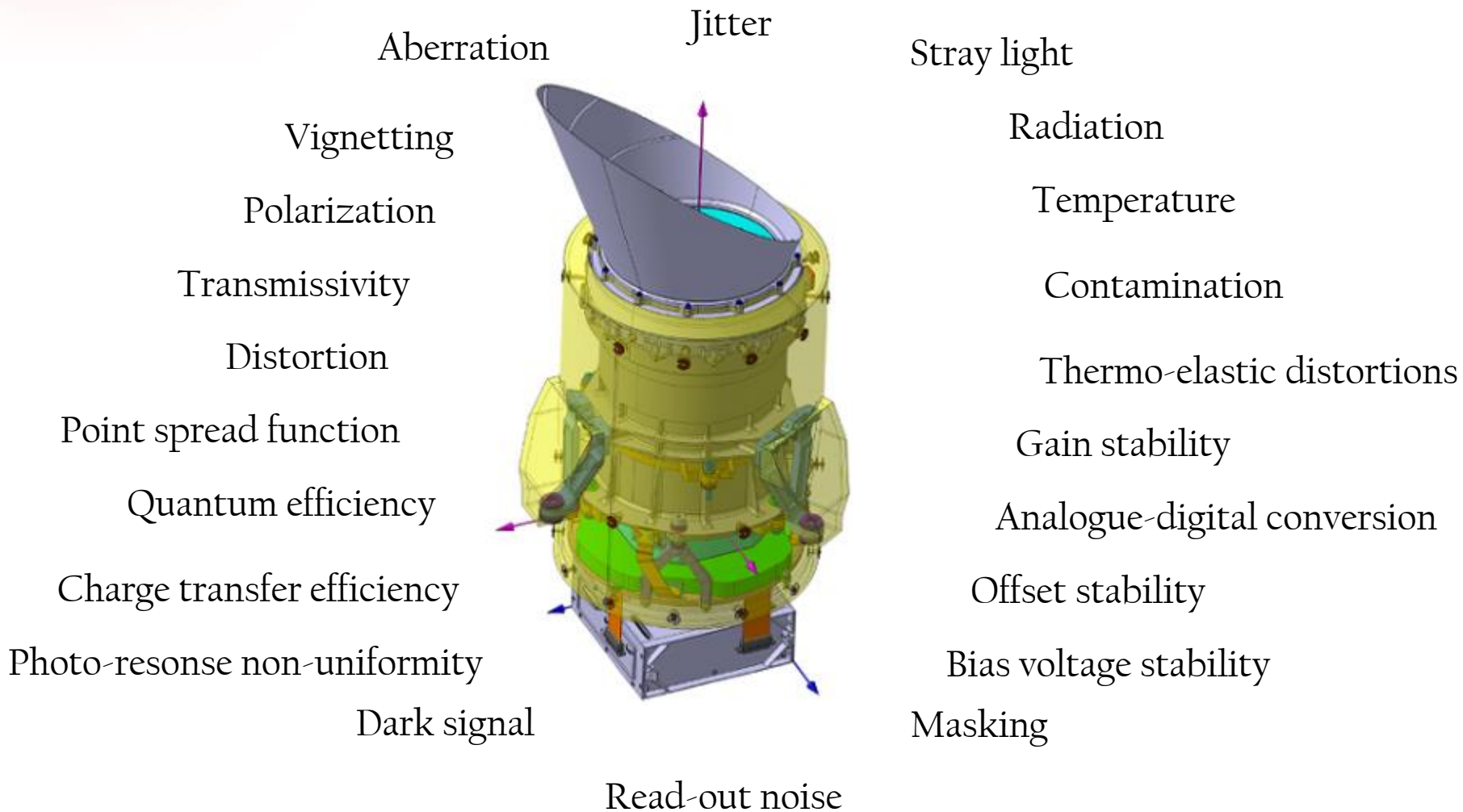


Instrument performance

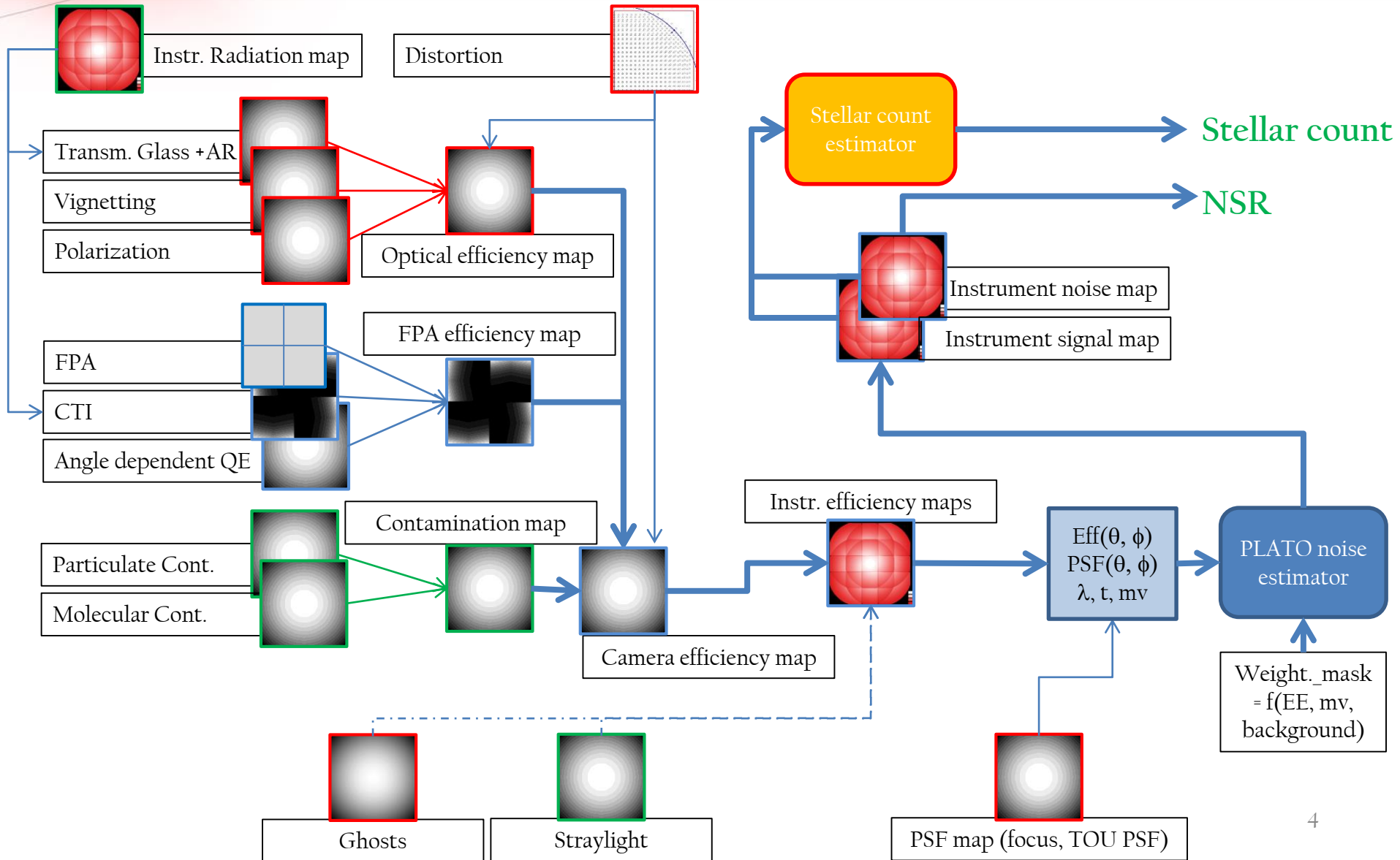
- Thank you to: Alan, Bart, Carsten, Dave, Demetrio, Denis, Gisbert, Jason, Joris, Juan, Jürgen, Martin, Mattheo, Matthias, Peter, Reza, Stefanie 1, Stefanie 2, Steve, Thibaut, Tomasz, Valerio, Valery ... from the entire team
- Motivation
 - NSR is PLATO's key performance parameter
 - NSR estimation is needed for
 - requirement definition and justification (PURD, TRD, URD, ...)
 - Sensitivity analysis
 - Optimization, mitigation and trade-off analysis
 - Input and cross validation to other simulation tools, e.g. PLATOSim
 - Input for data processing chains
- How?
 - Physical models
 - Spatially distributed maps



Performance impactors



Spatially resolved maps



Simulations

- Models

- Photons from target star to photoelectrons to digital values
- Photons from elsewhere
- Several other noise sources

- Software

- IDL program Noise estimator 6.8
- Two modes
 - Single pixel (run time < 1s)
 - Map (~1h per star field)

- Parameters

- See https://teamsites-extranet.dlr.de/pf/PLATO-Office/issue_1.4
- PPT -> Simulation parameter -> issue_1.4

```
PLATO_02_R_EOI_1.4.txt - Editor
Datei Bearbeiten Format Ansicht ?
-----
# PLATO
# system parameter table
# normal cameras, end of life, required values
# author: Anko Boerner, DLR
# date: 02.03.2018
-----
# Science
-----
000, NSR asteroseismology, ppm in 1 hr, 34
010, NSR planet transit detection, ppm in 1 hr, 80
-----
# Target
-----
020, reference star V magnitude, 0
030, reference star effective temperature, K, 6000
040, reference star (spectral type and luminosity class), text, G0V
050, reference star spectrum, text, Planck
060, reference photometric flux, w cm-2 micron-1, 3.6182e-12
070, reference photometric wavelength, nm, 550
080, target star (V magnitude), 11
090, contaminants and background model, text, simulated
100, reference star catalogue, text, PPMXL
110, southern long pointing field (SPF) galactic long, deg, 253
120, southern long pointing field (SPF) galactic lat, deg, -30
130, northern long pointing field (NPF) galactic long, deg, 65
140, northern long pointing field (NPF) galactic lat, deg, 30
150, incoming flux from target star, file, input_photon_density_m11.txt,
160, background signal, pe-/pix/s, 100
170, contaminant signal, -, 0.01
-----
# S/C
-----
200, jitter requirement 20uHz to 0.4Hz, arcsec/sqrt(Hz), 0.23
210, jitter requirement 20uHz to 3uHz, arcsec/sqrt(Hz), 0.23 to 21.4
220, jitter requirement > 0.4Hz, arcsec/sqrt(Hz), TBD
230, LOS jitter noise, file, jitter_R_17-09-22.txt,
235, S/C stray light, pe-/pix/s, 2
-----
# Instrument
-----
237, time frames of interest, file, time_frames_required_17-09-22.txt,
240, wavelength range for simulation start, nm, 500
250, wavelength range for simulation end, nm, 1000
260, wavelength range for simulation step, nm, 50
270, number of cameras, -, 22
280, angle between telescope groups w.r.t. instrument LOS, ', 9.2
285, CCD butting distance, mm, 2.6
290, exposure time, s, 21
300, cadence, s, 25
305, TCS frequency, Hz, 0.08
310, temperature measurement accuracy, file, T_meas_R_18-02-16.txt,
320, temperature stability TOU, file, T_stab_TOU_R_18-02-16.txt,
330, temperature stability CCD, file, T_stab_CCD_R_18-02-16.txt,
340, temperature stability FEE, file, T_stab_FEE_R_18-02-16.txt,
350, mean photon flux, e-/pix/s, 45
360, CCD bias stability, ppm/mv, 77
370, voltage measurement accuracy, mv, 0.2
380, clock stability, ppm in 14 h, 5
390, EMC noise, ppm, 40.
395, non ionizing radiation of CCDs, p-/cm2, 1.73e9
-----
# Optics
-----
400, effective pupil size diameter, cm, 12
410, effective aperture, m2, 0.01131
420, focal length, m, 0.24752
430, field of view 2D, deg2, 1037
432, overall optical FOV, deg2, 1120
434, corrected optical FOV (without CCD gaps), deg2, 1082
436, effective corrected optical FOV (including CCD gaps), deg2, 1037
438, field of view 1D, deg, 37.7
440, plate scale, arcsec, 15
445, distortion, -, 0.0384
450, mask for energy enclosed n by n, pixel, 1.0
460, energy enclosed in masked pixels, -, 0.32
480, transmission optics, file, optics_transmission_1.2pc_eoi_cutoff.txt,
490, transmission filters, file, filter_transmission_cutoff_eoi_v01.txt,
495, particulate contamination, ppm, 500*13+5000+650
```



Use cases

- For two star catalogues: PIC 0.2 and PIC 0.3
- Use cases
 - Standard
 - Beginning of life, required
 - End of life, required
 - Beginning of life, typical (e.g. CCD's QE can be 3% better than required)
 - End of life, typical
 - Padua cases
 - Changed exposure times
 - NPF + SPF
 - Evolution over life time
 - Extra
 - Changed divergence angle: $9.2^\circ \dots 0^\circ$

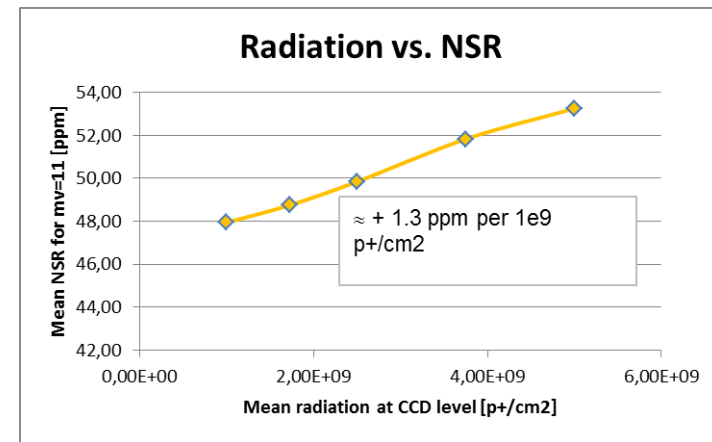
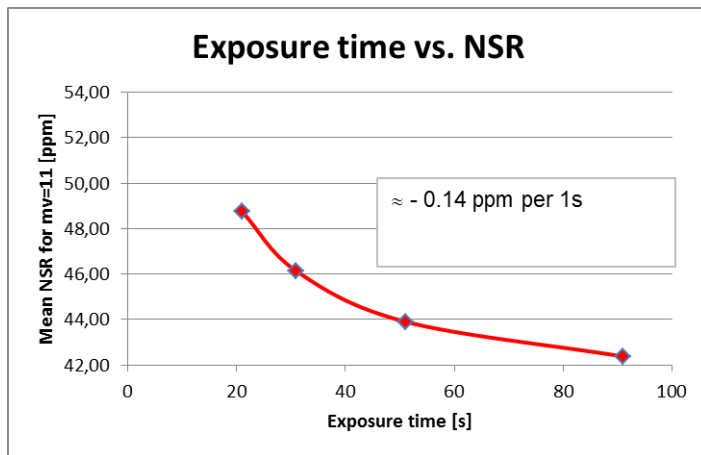
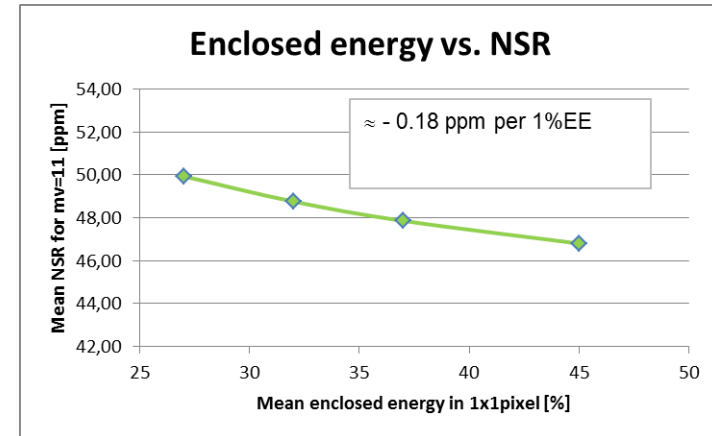
Results

- Enriched star catalogues contain NSR (random and random + systematic in 1hr) per star on instrument level (considering all cameras which see this star)

```
NA_backup\HD_D\Projects\PLATO\Engineering\Noise\Simulations_Support_and_Results\13_starcatalogue\padua_02\log_18-05-07\pic0_3_PLATO_Padua_001.txt - Notepad++
File Edit Search View Encoding Language Settings Tools Macro Run Plugins Window ?
pic0_3_PLATO_Padua_001.txt x
1 # PLATO
2 # -----
3 # 1. Column = Magnitude [-]
4 # 2. Column = Longitude in S/C coordinate system [deg]
5 # 3. Column = Latitude in S/C coordinate system [deg]
6 # 4. Column = random NSR [ppm]
7 # 5. Column = random + systematic NSR [ppm]
8 # 6. Column = number of cameras seeing the star
9 # 7. Column = number of cameras getting saturated
10 # 8. Column = sigma of cameras PSF
11 # 9. Column = DEBUG_p_target
12 # 10. Column = DEBUG_opt_trans
13 # 11. Column = DEBUG_add_efficiency
14 # 12. Column = DEBUG_overall_eff_wo_cti
15 # 13. Column = DEBUG_e_target
16 # 14. Column = DEBUG_overall_eff
17 # 15. Column = DEBUG_eff_CTI
18 # 16. Column = DEBUG_e_target_cti
19 # 17. Column = DEBUG_signal_spread
20 # 18. Column = DEBUG_external_eff
21 # 19. Column = star ID
22 # -----
23 # Simulation coverage [%]:      100.000
24 # -----
25 10.27 -12.97 24.21 7.369976e+001 7.428741e+001 5 0 0.640 920341. 0.798 0.838 0.366 90881. 0.341 0.933 84780. 0.270 0.583 3597046
26 10.28 -13.07 24.17 7.380794e+001 7.439474e+001 5 0 0.640 914006. 0.798 0.838 0.366 90441. 0.342 0.933 84404. 0.270 0.584 3597486
27 10.28 -14.85 23.56 7.317720e+001 7.376900e+001 5 0 0.640 918310. 0.798 0.838 0.365 90454. 0.344 0.943 85266. 0.270 0.581 3599886
28 9.99 -13.01 24.04 6.275879e+001 6.344756e+001 5 0 0.640 1190335. 0.798 0.838 0.368 118424. 0.348 0.943 111690. 0.270 0.587 3602944
29 10.80 -15.83 23.15 9.819982e+001 9.864218e+001 5 0 0.640 566121. 0.798 0.838 0.363 55539. 0.338 0.930 51660. 0.270 0.579 3603278
30 10.47 -12.13 24.25 8.188943e+001 8.241890e+001 5 0 0.640 769394. 0.798 0.838 0.369 76630. 0.340 0.921 70557. 0.270 0.588 3604560
31 10.00 -15.53 23.21 6.294257e+001 6.362936e+001 5 0 0.640 1182032. 0.798 0.838 0.365 116378. 0.348 0.954 111042. 0.270 0.581 3604898
32 10.94 -15.72 23.14 1.072480e+002 1.076534e+002 5 0 0.640 498413. 0.798 0.838 0.364 49036. 0.337 0.924 45309. 0.270 0.581 3605308
33 10.97 -14.30 23.56 1.080995e+002 1.085017e+002 5 0 0.640 485276. 0.798 0.838 0.368 48277. 0.336 0.912 44007. 0.270 0.587 3606502
34 10.01 -16.85 22.70 6.325337e+001 6.393683e+001 5 0 0.640 1175626. 0.798 0.838 0.361 114665. 0.347 0.960 110079. 0.270 0.576 3607278
35 10.65 -11.76 24.28 9.093893e+001 9.141621e+001 5 0 0.640 652693. 0.798 0.838 0.370 65213. 0.337 0.910 59361. 0.270 0.590 3607352
36 10.97 -12.07 24.17 1.086988e+002 1.090989e+002 5 0 0.640 486798. 0.798 0.838 0.372 48884. 0.333 0.896 43777. 0.270 0.593 3608151
37 10.78 -12.58 24.02 9.729052e+001 9.773695e+001 5 0 0.640 575373. 0.798 0.838 0.372 57771. 0.338 0.909 52503. 0.270 0.593 3608270
38 10.50 -14.57 23.40 8.386893e+001 8.438599e+001 5 0 0.640 750912. 0.798 0.838 0.369 74835. 0.345 0.935 69949. 0.270 0.588 3609210
39 10.78 -13.04 23.86 9.675403e+001 9.720291e+001 5 0 0.640 579094. 0.798 0.838 0.372 58108. 0.339 0.912 53018. 0.270 0.592 3609251
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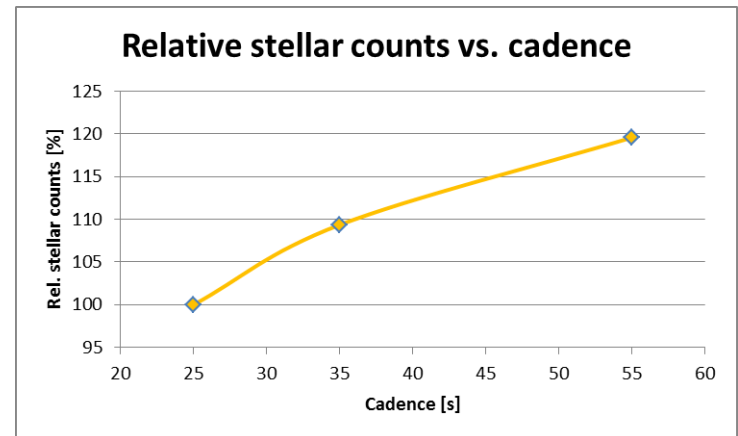
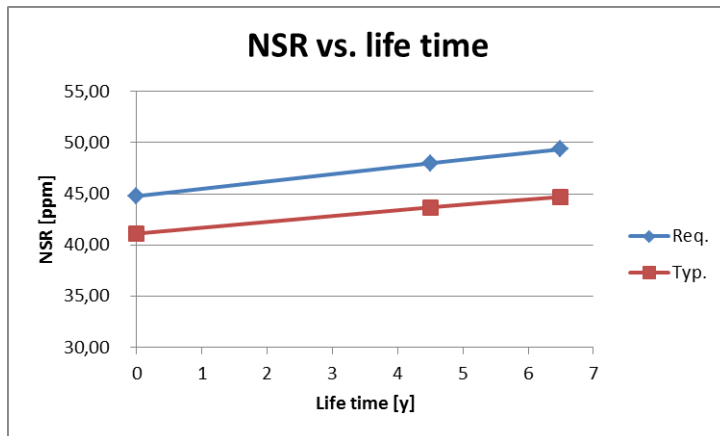
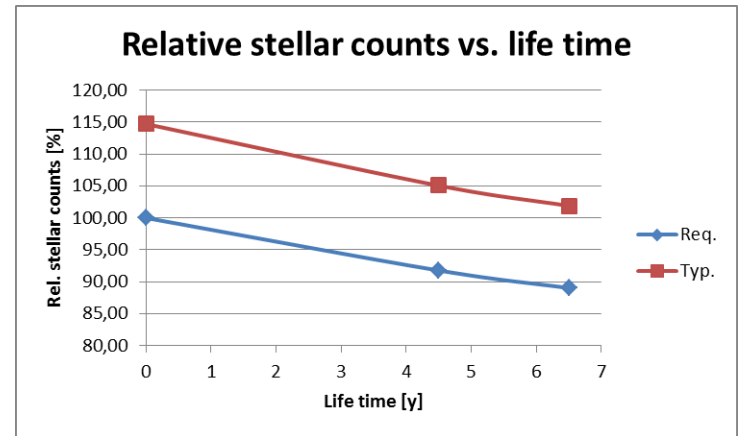
Results

- Sensitivity analysis (single pixel mode)
 - Fast but average value only
 - To understand dependencies and impacts of single parameters to NSR



Results

- Sensitivity analysis (map mode)
 - Slow but accurate
 - Each star is assigned to NSR
 - Stellar counts can be predicted





Next steps

- Support of several requirement activities
- Support of signal/ noise modelling activities, e.g. transient simulations, FCam science
- Stellar counts, planet yields can be estimated based on the enriched star catalogues, update of the star catalogue expected (PIC0.4)